

# The Degenerate MSSM

Jesse Thaler

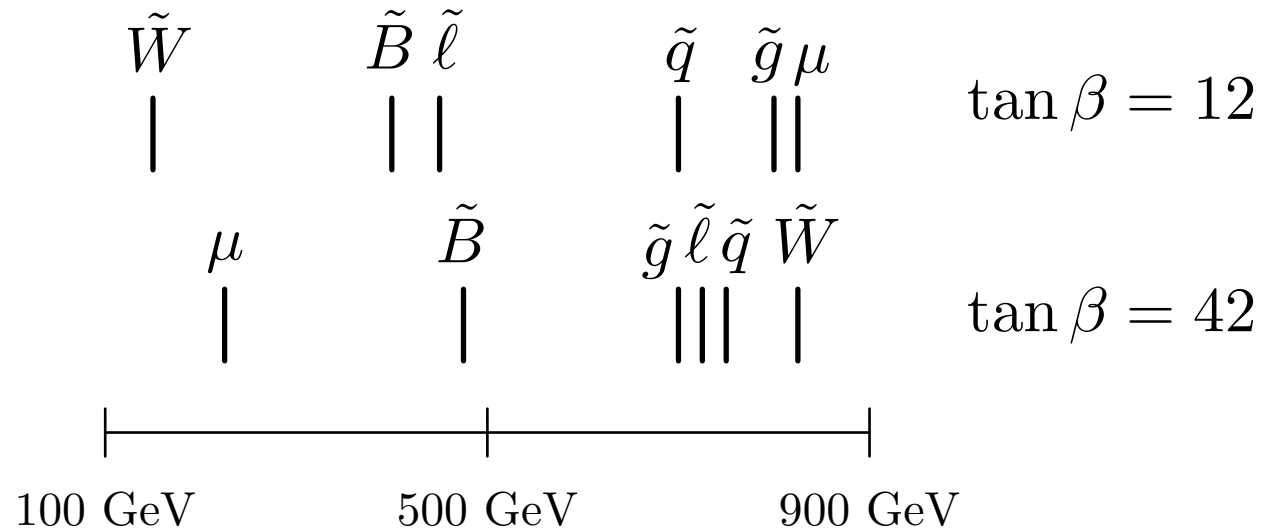
with N. Arkani-Hamed, G. Kane, and L.-T. Wang

# One Year at the LHC

$10 \text{ fb}^{-1}$

Experimental Data  $\longrightarrow$  Theoretical Models?

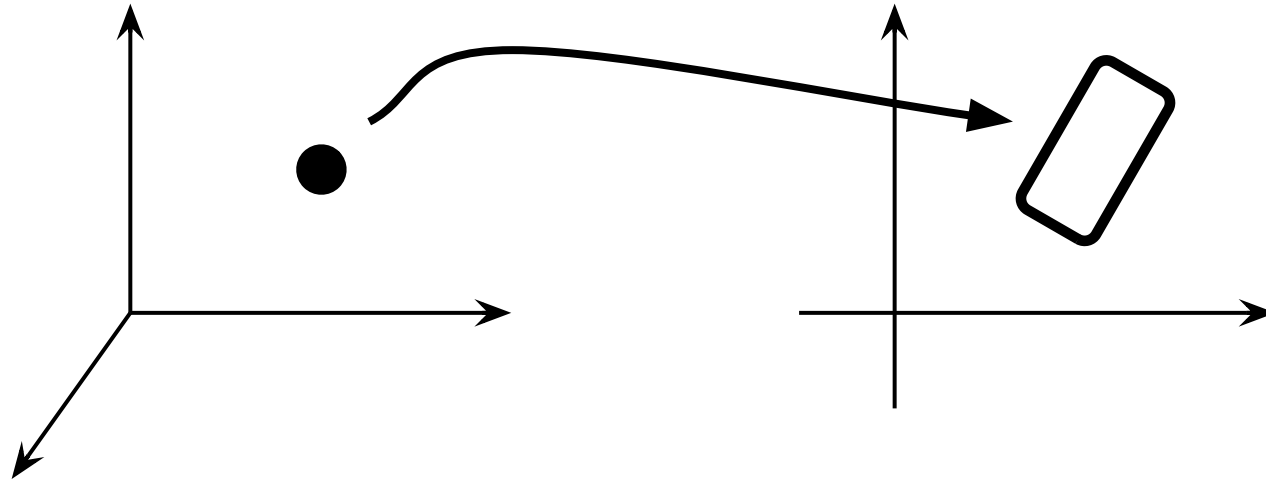
If it's the  
MSSM...



# Standard Method

Parameter Space

Signature Space



TDR, Benchmark studies, LEP/Tevatron bounds, etc.

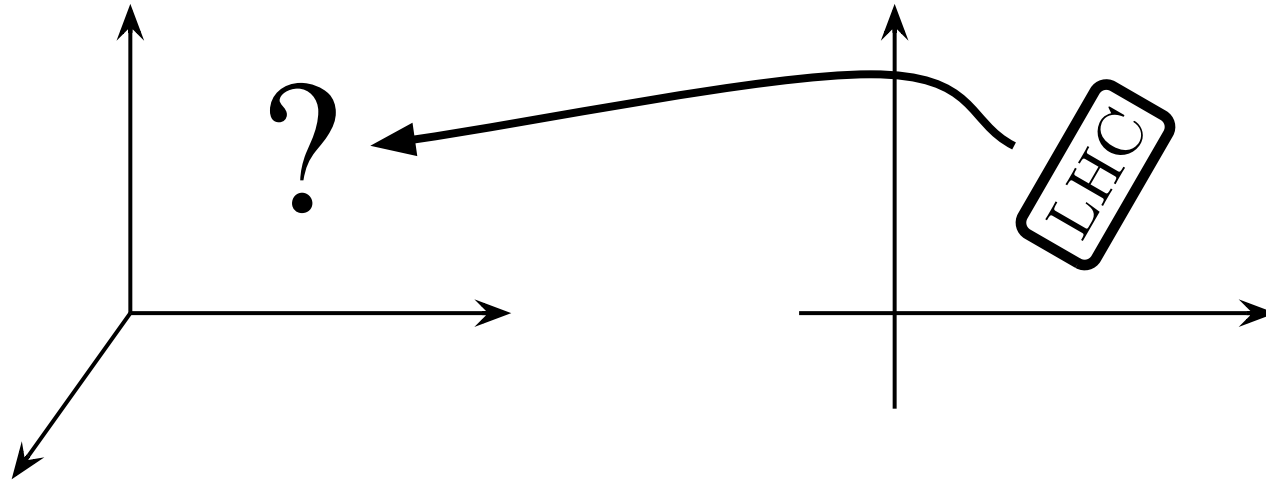
Reduced parameter set (mSUGRA, AMSB, GMSB):

Experimental Data  $\longrightarrow$  Precision Measurements

# The Inverse Problem

Parameter Space

Signature Space

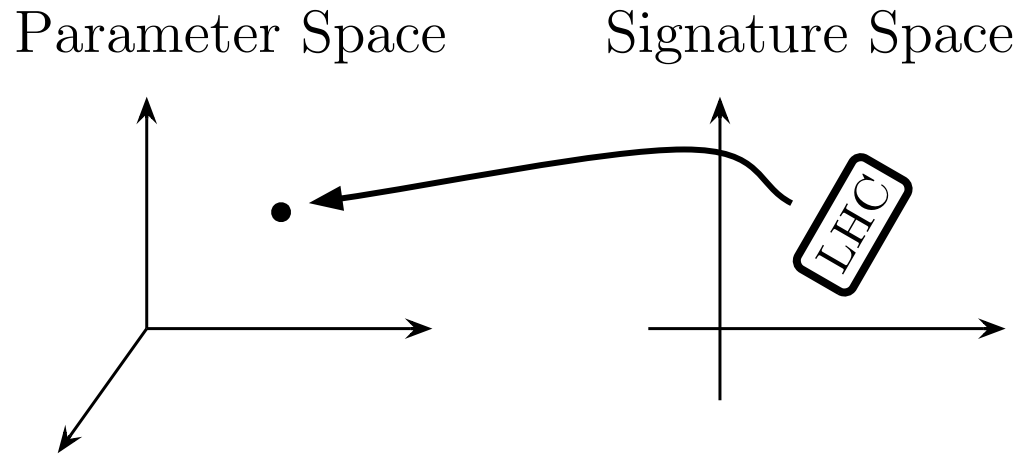


Much more interesting!

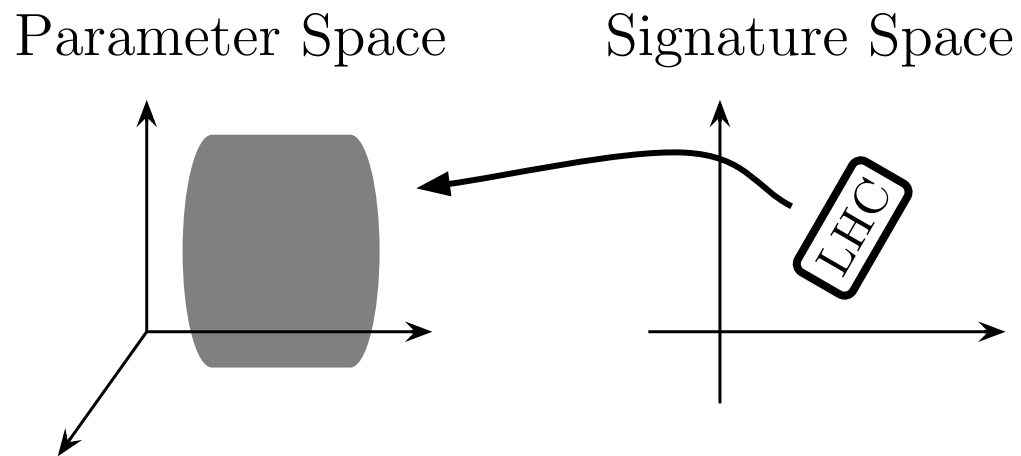
Data  $\xrightarrow{?}$  gaugino unification, dark matter, ...

Much more important! (500 GeV ILC in 10 years?)

# Best of all Possible Worlds



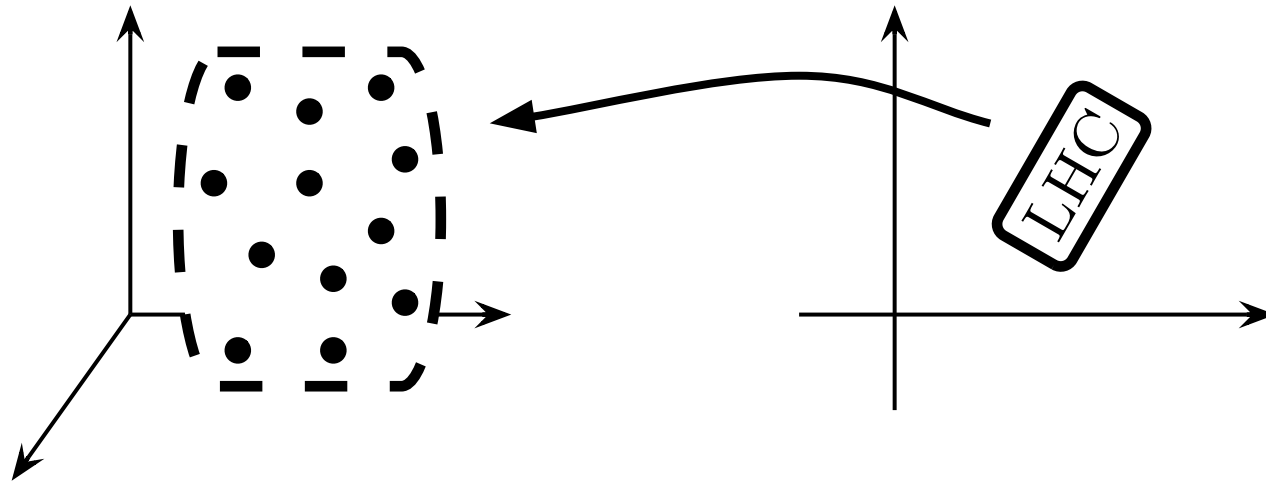
# Worst of all Possible Worlds



# The Real MSSM

Parameter Space

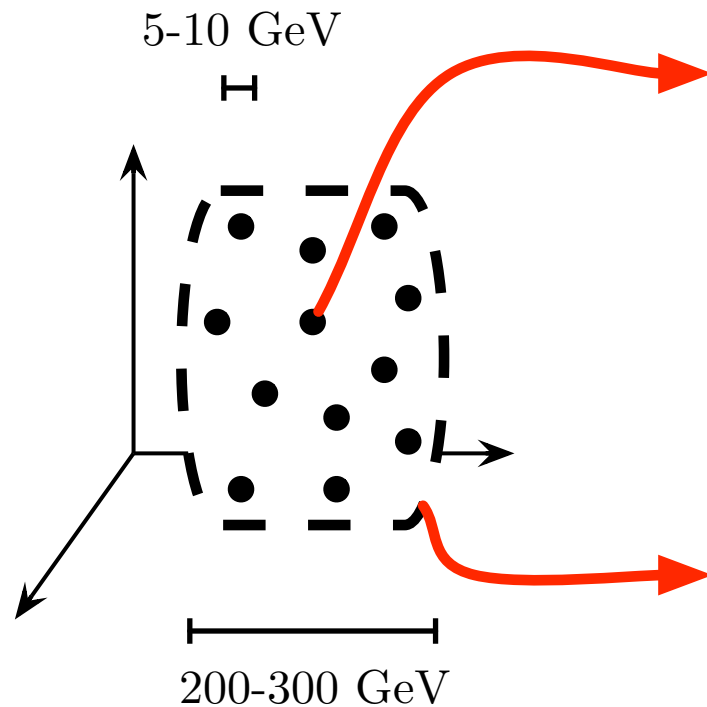
Signature Space



## Degeneracies!

Many small footprints in a large overall region.

# The Degenerate MSSM

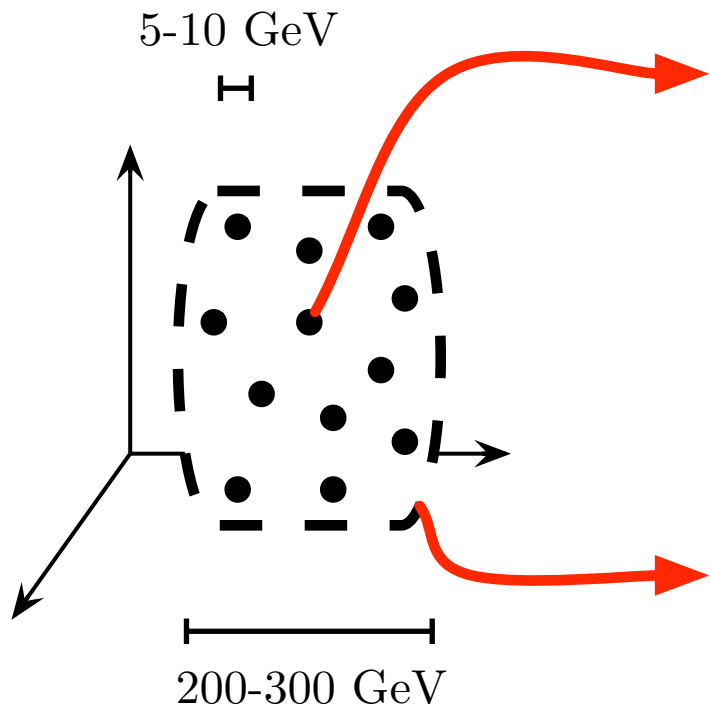


If LHC data is consistent with mSUGRA, you will measure mSUGRA parameters very accurately...

...but there at least  $\mathcal{O}(100)$  other (well-motivated?) MSSMs consistent with the same experimental data.

200 GeV can change gaugino unification, LSP identity, etc.

# On The Snowmass Slopes

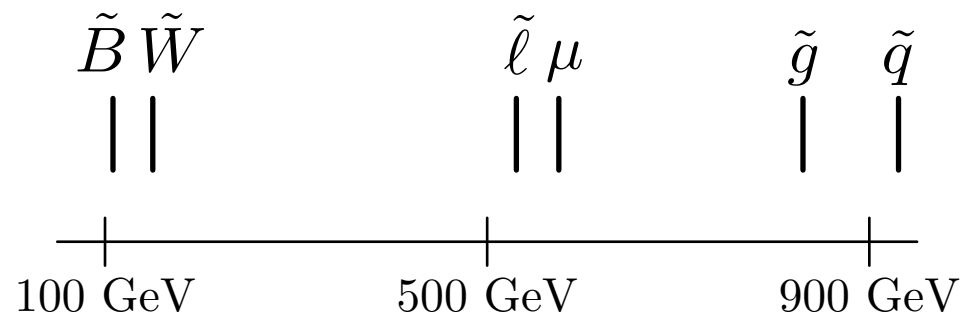


Slope 1a: (Allanach, et al, 2002.)

$$m_{1/2} = 425 \text{ GeV}$$

$$m_0 = -A_0 = 170 \text{ GeV}$$

$$\mu > 0, \quad \tan \beta = 10$$



$$\tan \beta = 46$$

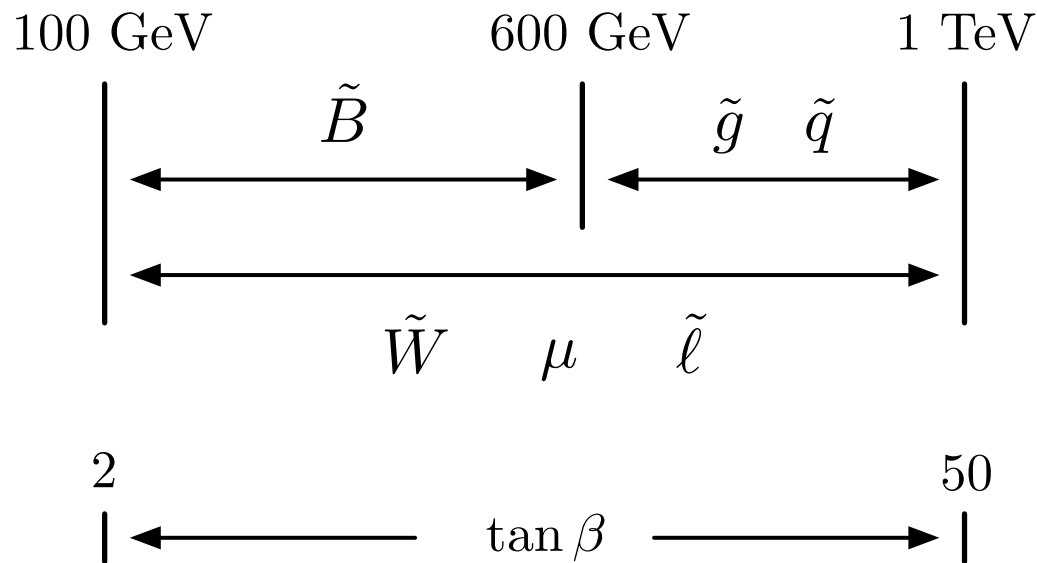
$$\chi_{\text{sig}}^2 = 5$$



# Strategy

## 50,000 MSSMs

Nowhere near full coverage, no flavor dependence.



No slepton LSP. Everything else at PYTHIA default values.

# Inclusive Signatures

$$2 \text{ jets} + \cancel{E}_T + \left\{ \begin{array}{l} 0\ell \\ 1\ell \\ \text{OS} \\ \text{SS} \\ 3\ell \\ 4^+\ell \end{array} \right. \begin{array}{l} \\ \\ \text{jet multiplicity} \\ b\text{-jet multiplicity} \\ \\ \\ \end{array} \begin{array}{l} \langle \cancel{E}_T \rangle \\ \langle E_T^{\text{lep}} \rangle \\ \langle E_T^{\text{jet}} \rangle \end{array}$$

(Baer, Chen, Paige, Tata, 1995.)

Total of 22 (very broadbrush) signatures.

Real LHC observables.

Standard TDR-like cuts. No SM background.

Assume 10% error.

What can you do with 50,000 MSSMs?

$(50000)^2$  Pairs!



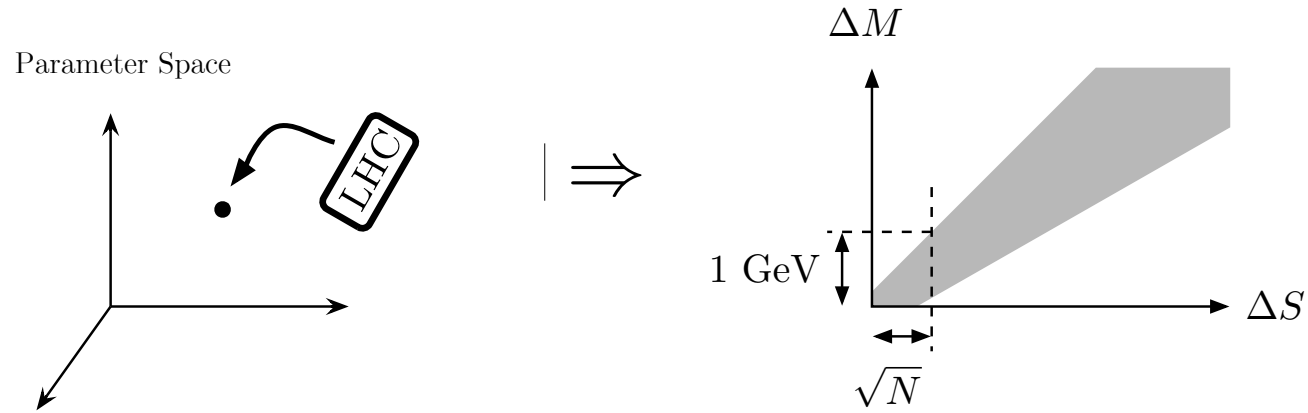
Statistics!

For each pair, signature vs. parameter distance.

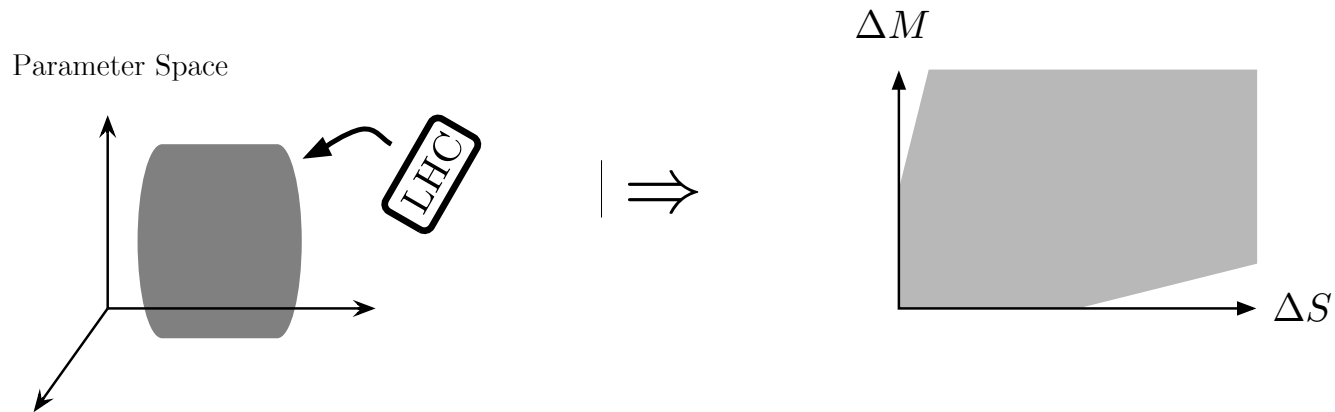
$$\Delta S = \chi_{\text{sig}}^2 = \sum_i \left( \frac{\Delta s_i}{\sigma_i} \right)^2$$

$$\Delta M = \text{“}\chi_{\text{para}}^2\text{”} = \sum_j \left( \frac{\Delta m_j}{\text{“}\sigma_j\text{”}} \right)^2$$

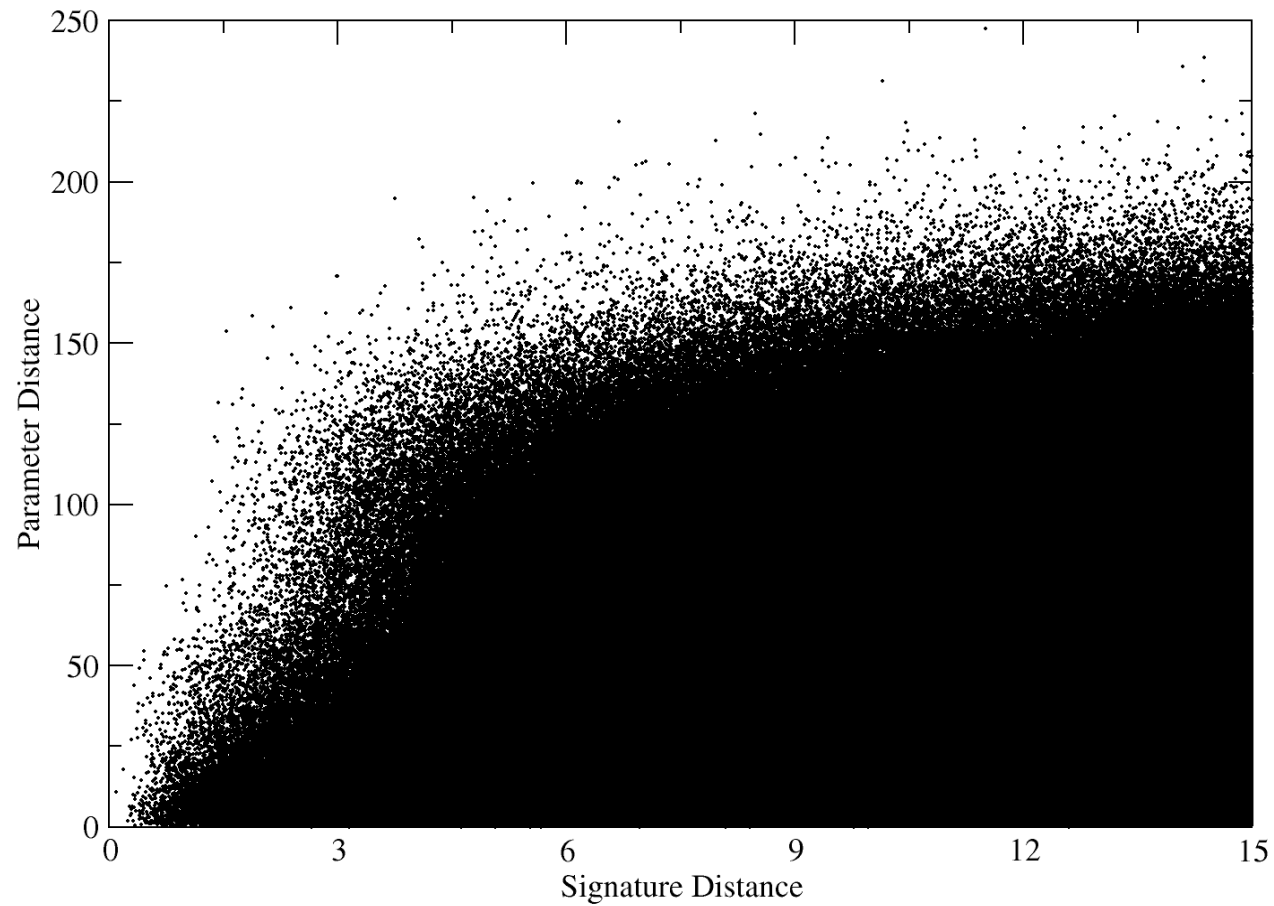
# Best of all Possible Worlds



# Worst of all Possible Worlds

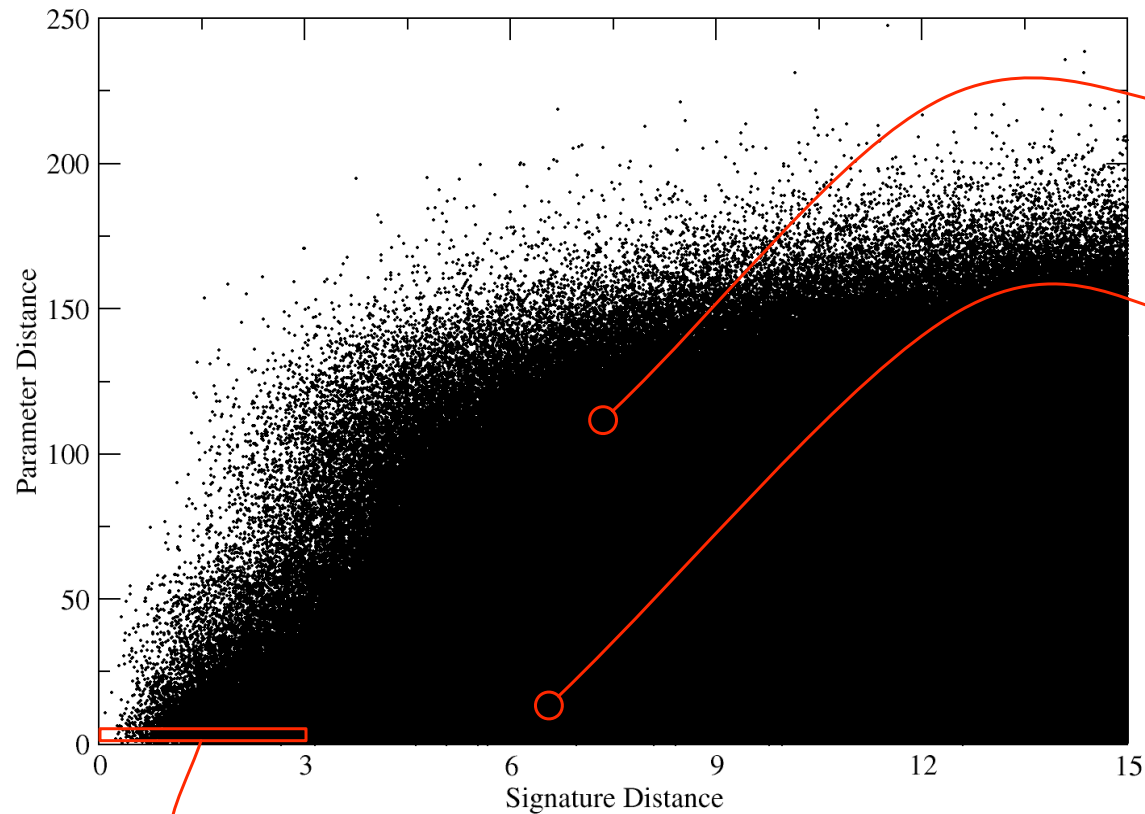


# The Real MSSM



Shown:  $4.9 \times 10^6$  of  $2.6 \times 10^9$  pairs.

# The Real MSSM



MSSM 2005 vs.  
MSSM 22725

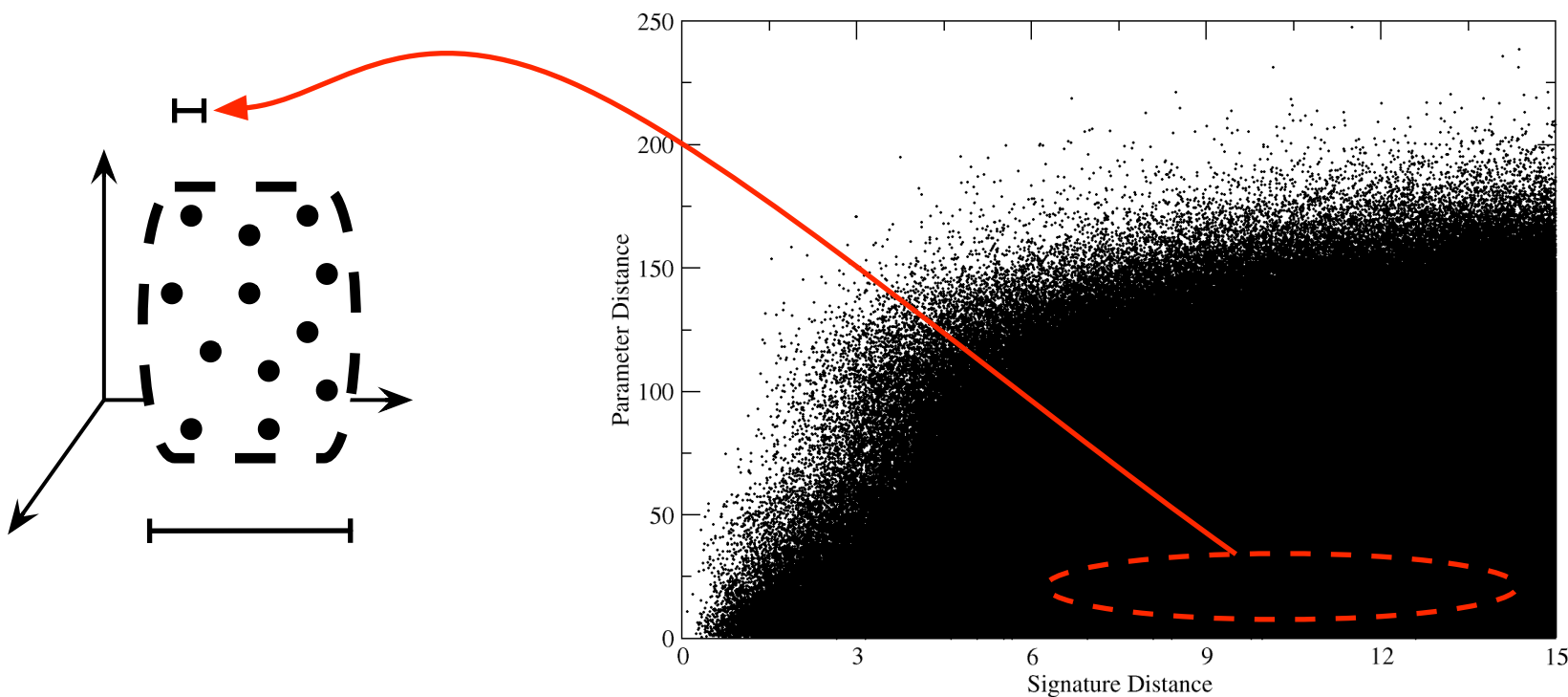
MSSM 2005 vs.  
MSSM 47553

LHC Success Region

$$\left\{ \begin{array}{l} \chi_{\text{para}}^2 < .75 \leftrightarrow 5\% \text{ error} \\ \chi_{\text{sig}}^2 < 3.0 \leftrightarrow 1.5 \times \sqrt{N} \text{ error} \end{array} \right.$$

# Two Surprises: Cliffs

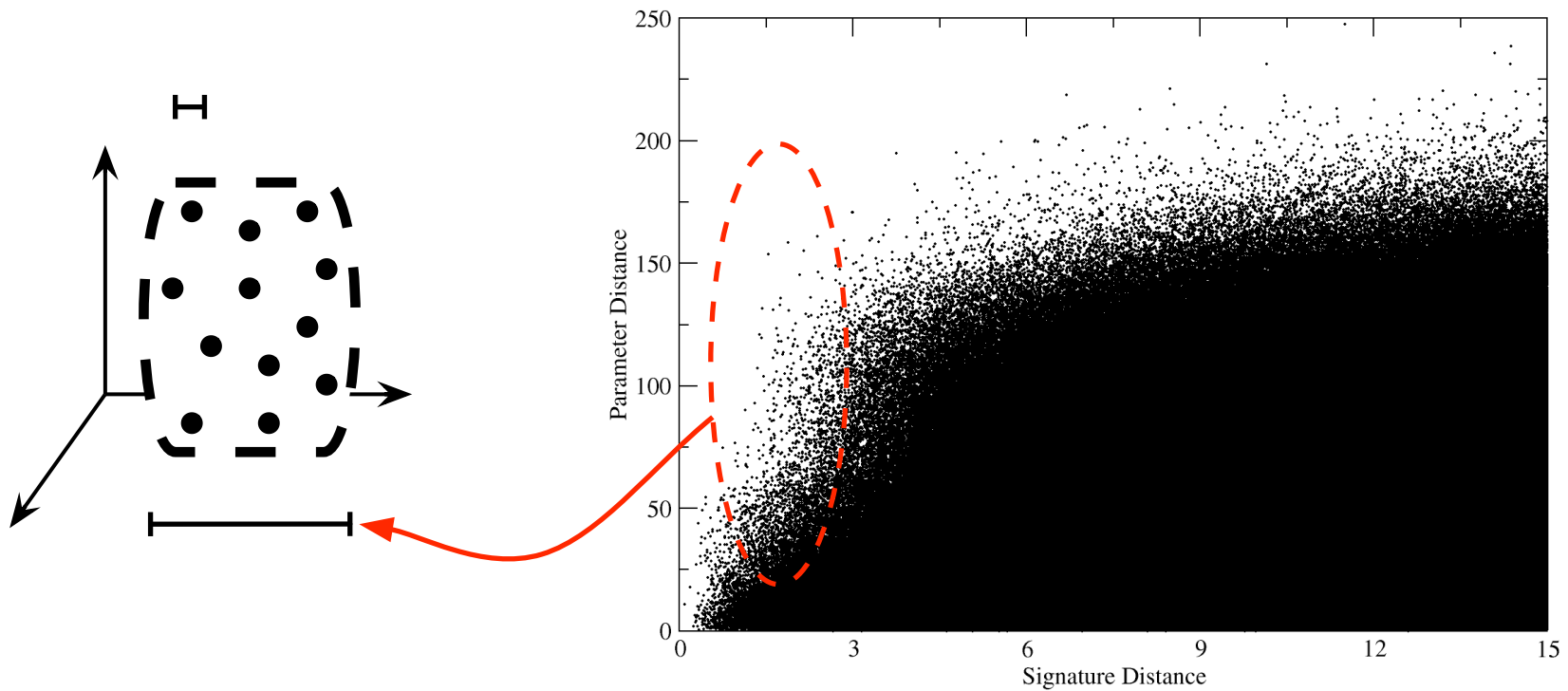
Close in parameter space, far away in signature space.



Evidence for small individual footprint size.

# Two Surprises: Degeneracies

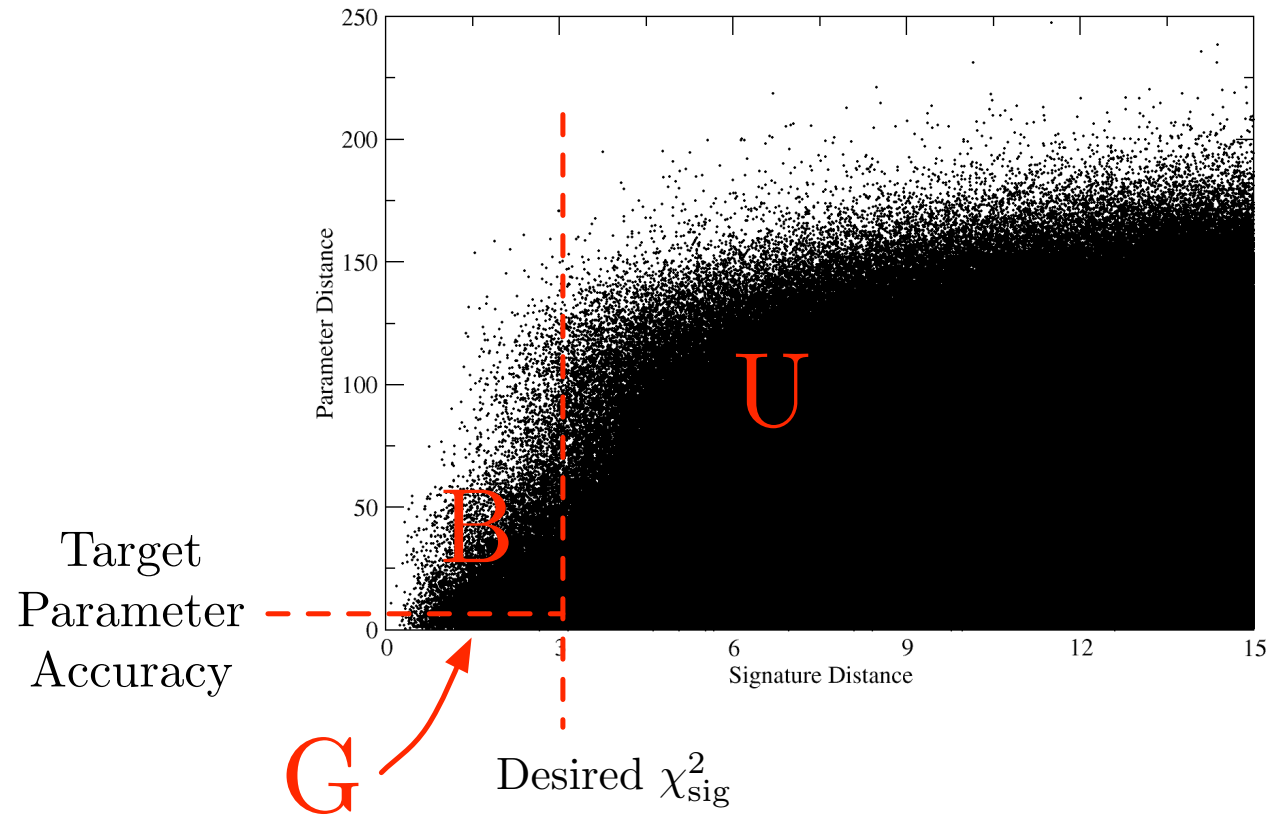
Close in signature space, far away in parameter space.



Evidence for large number of footprints.

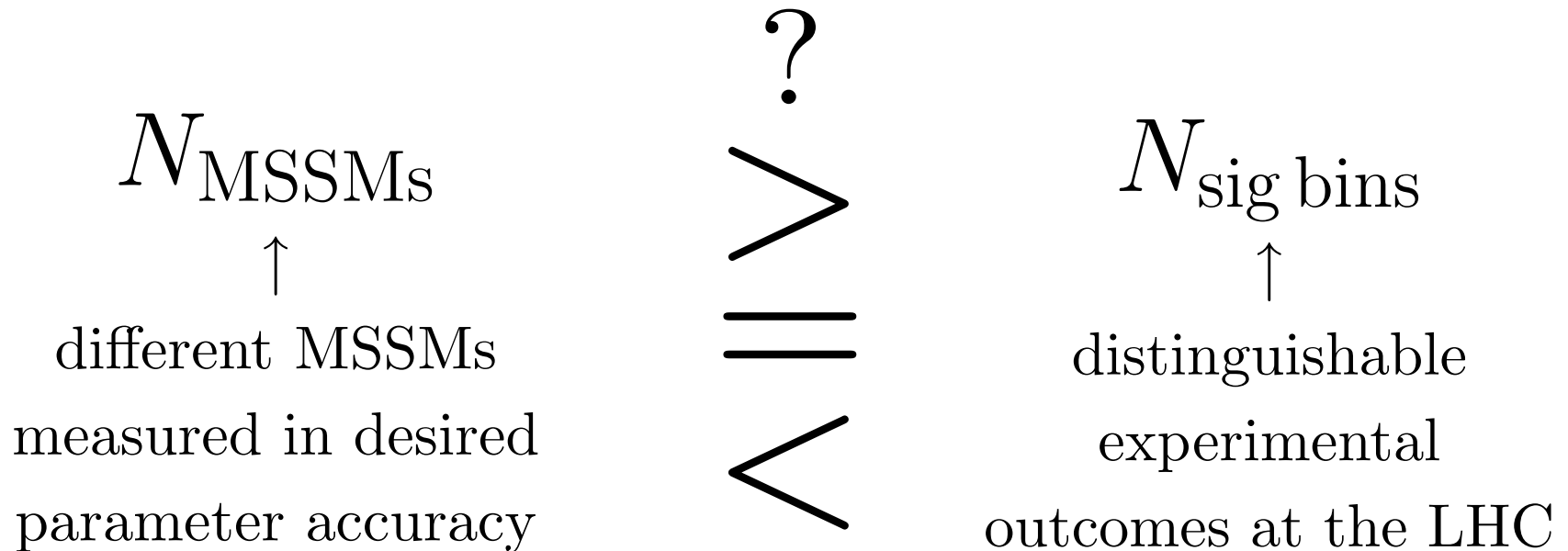


# Counting Degeneracies



$$\langle d \rangle = \frac{G + B}{G} \sim 100 (!) \quad \chi_{\text{sig}}^2 < 3 \quad \Delta M \lesssim 5\%$$

# Why So Many Degeneracies?



Pigeon hole principle!

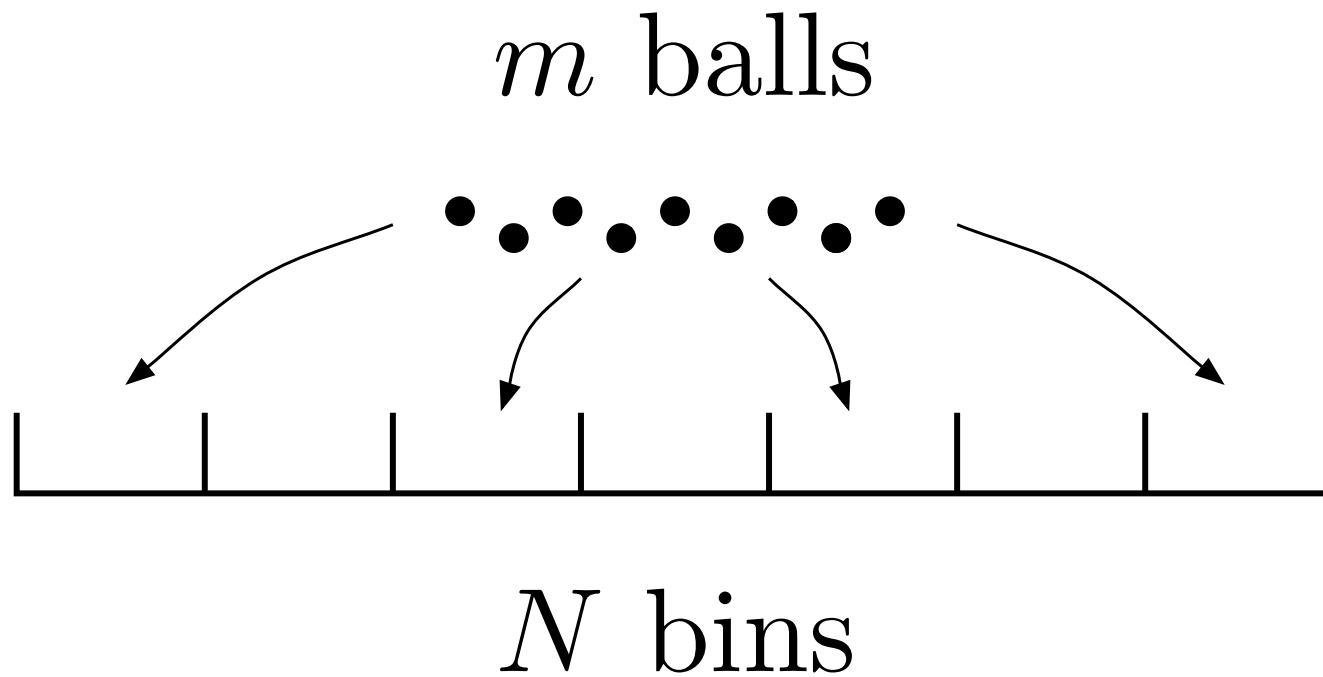
5% bins in parameter space (modulo decoupling):

$$N_{\text{MSSMs}} \sim 10^7$$

Simple way to count signature bins for given  $\chi_{\text{sig}}^2$ :

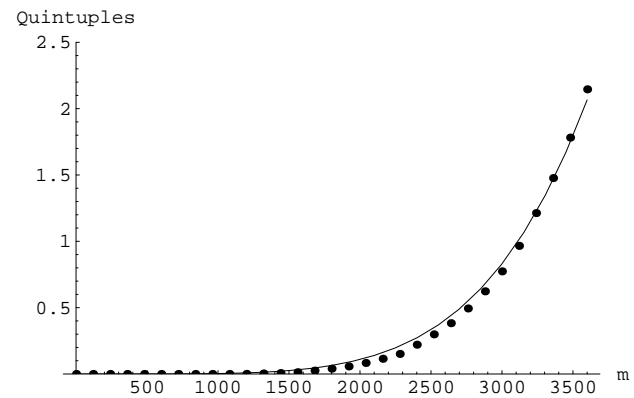
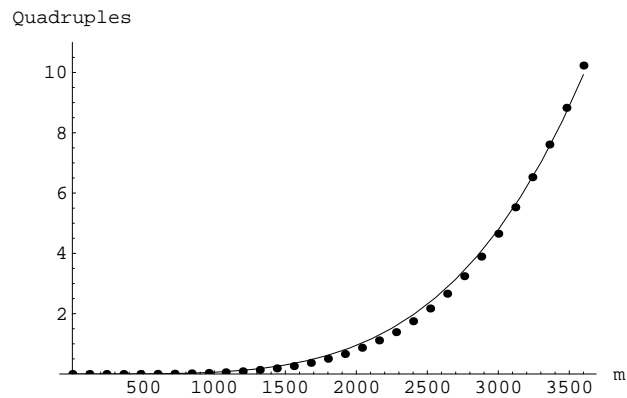
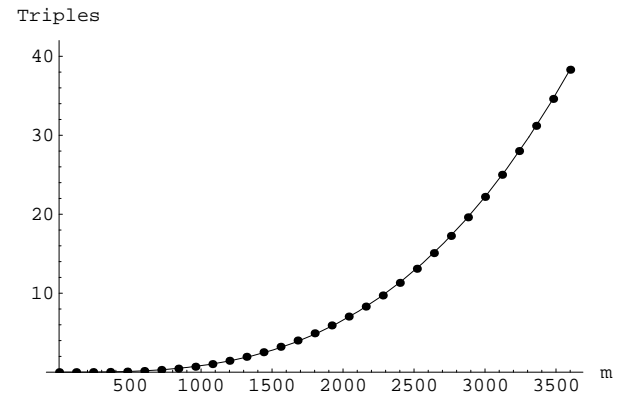
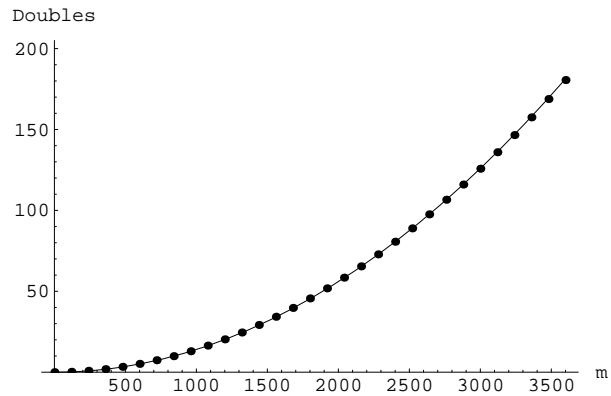
$$N_{\text{sig bins}} = \left\{ \begin{array}{ll} 310,000 & \chi_{\text{sig}}^2 < 2 \\ 83,000 & \chi_{\text{sig}}^2 < 3 \\ 5,000 & \chi_{\text{sig}}^2 < 6 \\ 1,600 & \chi_{\text{sig}}^2 < 9 \\ 800 & \chi_{\text{sig}}^2 < 12 \\ 400 & \chi_{\text{sig}}^2 < 15 \end{array} \right.$$

# Statistics



$$\begin{array}{c} \bullet \quad \bullet \\ \swarrow \quad \searrow \\ \boxed{\phantom{\bullet \quad \bullet}} \end{array} = \frac{1}{2!} \frac{m^2}{N} \qquad \begin{array}{c} \bullet \\ \downarrow \\ \bullet \quad \bullet \\ \swarrow \quad \searrow \\ \boxed{\phantom{\bullet \quad \bullet}} \end{array} = \frac{1}{3!} \frac{m^3}{N^2}$$

# Statistics of the MSSM



$$N_{\text{sig bins}} \sim 83,000$$

$$\chi_{\text{sig}}^2 \lesssim 3$$

# Reconciling Two Different Attitudes

SUSY at LHC is easy!

If a constrained model is consistent with LHC data, then inclusive signatures can yield precision measurements at  $10 \text{ fb}^{-1}$ .

SUSY at LHC is hard!

There are  $\mathcal{O}(100)$  different models that also match the data, and there is no systematic way to find them. Do they come from nice UV theories?

# Reality Checks

Qualitatively the same story if we:

Use  $\sqrt{N}$  error.

Enforce  $> 10000$  SUSY events.

Artificially include dilepton edges/endpoints.

Including flavor information:

7 Parameters  $\rightarrow$  14 Parameters

$10^7 \rightarrow 10^{14}$  MSSMs

$10^5 \rightarrow 10^6$  Signature Bins

Even more degeneracies!

# Lessons from the Degenerate MSSM

LHC Signature Space



MSSM Parameter Space

To nail the MSSM at the LHC, we must drastically increase the number of independent inclusive signatures.

Easy to check if a new inclusive works: calculate  $\langle d \rangle$ .

MSSM vs. NMSSM vs. UED vs.  $T$ -parity LH ?